Non-paraxial Talbot Effect for Building Compact Spectrometers

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Abstract: We explore the possibility of utilizing mid-field Talbot effect under non-paraxial diffraction for building compact spectrometers. Our experiment demonstrates nanometer resolution with a bandwidth over at least 100 nm for a Talbot spectrometer built using a standard 1-D transmission grating and a commercial CMOS imager. The effect of angular spread at the grating surface on spectral resolution is also investigated theoretically.

OCIS codes: 070.6760, 120.6200, 100.3190

1. Introduction

Computational spectrometers, either fiber/waveguide [1,2] or free space-based [3,4], emerge as interesting candidates for modern spectroscopy applications. As opposed to conventional grating-based spectrometers, these computational spectrometers do not require explicit orthogonal imaging in the far-field for wavelength distinction, but rather employ a non-orthogonal (but still invertible) wavelength dispersion mechanism with subsequent machine learning-like calibration and reconstruction processes. Examples of such dispersion mechanisms include 3-D photonic crystals [3], multimode fibers [2], and pseudo-random diffractive optics [4] among others. Because of the elimination of explicit orthogonal imaging optics, a computational approach opens possibilities for spectrometry with unconventional form factors as well as enhanced sensitivity due to the multiplexing advantage. Here, with a similar design principle in mind, we explore the possibility of utilizing the mid-field Talbot effect [5] under non-paraxial conditions as the dispersion mechanism for building a compact spectrometer.





2. Theory

The Talbot self-images occur when a spatially periodic object such as a transmission grating is illuminated by spatially coherent light. The self-image replicates itself along the direction perpendicular to the grating plane (the *z* direction) in an alternating phase-reversing fashion as shown in Figure 1 (a). In order to utilize the Talbot effect to build compact spectrometers, we constrain the Talbot self-images to be within a few millimeters behind the grating by choosing the grating pitch to operation under non-paraxial diffraction for the wavelengths of interest. The periodicity of the self-images z_T as a function of the wavelength λ in this case can be characterized as $z_T(\lambda) = \lambda (1 - \sqrt{1 - \lambda^2/d^2})^{-1}$, where *d* is the grating period. Assuming in the general case of arbitrary diffraction efficiencies of A_0 , A_1 and A_{-1} for the three diffracted beams, the field intensity after the grating can be characterized as

$$I(x,z) \propto A_0^2 + A_1^2 + A_{-1}^2 + 2A_1A_{-1}\cos(4\pi x/d) + 2A_0(A_1 + A_{-1})\cos(2\pi x/d)\cos(2\pi z/z_T) + 2A_0(A_1 - A_{-1})\sin(2\pi x/d)\sin(2\pi z/z_T).$$

which is periodic in the *z* direction with the periodicity of z_T that determines the corresponding wavelength. A tilted CMOS imager is placed right behind the grating to sample the Talbot patterns across the *z* direction as shown in Figure 1 (b). The captured image can be subsequently transformed into spectrum of the source through standard Fourier transforms. The spectral resolution and maximum bandwidth without aliasing of the spectrometer can be determined by the total sampling length and spatial sampling period of the Talbot diffraction pattern, given by the projected sensor length and pixel size in the *z* direction due to the camera tilt θ_{tilt} . Assuming that the sensor length is *L* and the pixel size is z_{pix} , the resolution Δk_T and bandwidth $k_{T,max}$ in the spatial Fourier domain are determined as $\Delta k_T = 2\pi (L \sin \theta_{tilt})^{-1}$ and $k_{T,max} = \pi (z_{pix} \sin \theta_{tilt})^{-1}$, where $k_T = 2\pi / z_T$. This provides a simple design guideline for a resolution and bandwidth target.

3. Experiment

To verify our model, we characterized spectrometers with off-the-shelf 1D transmission gratings and commercial CMOS imagers. With a size constraint in mind, we chose CMOS imagers for smartphone and compact camera applications, where the pixel size is at most a few micrometers. Our grating has a grating pitch of $d = 1.035 \mu m$ and our imager has a pixel size of $z_{pix} = 1.12 \mu m$. The CMOS imager does not have any color filters, whereas the micro-lens array remains on top over the pixels. The glass protection window has been removed to prevent any spurious reflection. The imager has a dimension of 4208×3120 pixels (4.713 mm $\times 3.494$ mm), where the longer dimension is used to sample across the z direction to get a higher resolution. For characterizing our spectrometer, a tunable Ti:Sapphire laser in CW mode is coupled into a single mode fiber. The output of the fiber passes through a fiber collimation lens and a 10X beam expander. The collimated output beam is normally incident on the grating and the resulting Talbot pattern is recorded by the CMOS imager. For simplicity of characterization, the imager is placed to sample only the $+1^{st}$ and the 0th order diffracted beams.



Fig. 2. (a) Simulated and captured Talbot pattern captured by the CMOS imager. (b) Measured spectral resolution as a function of tilt angle compared with the theory. (c) Transformed spectral peaks obtained from experiment with a tunable Ti:Sapphire laser across 100 nm.

Simulated and captured images for the sampled Talbot pattern are shown in Figure 2 (a) for an imager tilt angle of 30°. The periodic pattern of the self-images can be seen clearly from both images and the experiment agrees well with the simulation. To further verify our model of the Talbot spectrometer, the peak resolution, characterized as the full width at half maximum (FWHM) of the transformed spectral peaks, at 830 nm is plotted as a function of θ_{det} in Figure 2 (b). A 10-row subsection of the full image is used for spectral reconstruction to avoid peak position variation across the image rows, possibly due to sampling distortion introduced by the presumed non-uniform displacement of the micro-lens array known as chief ray angel correction. As indicated in the figure, the resolution increases as the tilt angle increases as the theory predicts. At a tile angle of 30°, a resolution close to 1 nm (1.3 nm) is achieved. A consistent resolution overestimation from the theory is observed across all the tilt angles, which could be attributed

to peak broadening due to pixel cross-talk for non-normal light incidence on the imager or the imager chief ray angle correction mentioned earlier. A series of transformed spectra across 100 nm from 790 to 890 nm at a tilt angle of 30° is plotted in Figure 3 (c), showing the wavelength dependency of the transformed spectral peaks.

4. Discussions

Our experimental results suggest that the non-paraxial Talbot effect can be used as the dispersion mechanism for building a compact spectrometer with high resolution. A natural question that arises is how much spatial incoherence one can tolerate for a give resolution of such a spectrometer. We theoretically characterized the effective resolution with a given incidence angle spread over the polar and azimuthal directions on the grating for the Talbot spectrometer. Figure 3 (a) shows our simulation model, where the Talbot wave vector k_T arises due to the interference between the $+1^{st}$ and -1^{st} order diffracted beams with the 0th order diffracted beam. An oblique incidence beam results in a shift in the Talbot wave vector k_T compared to the normal incidence case, and the shift depends on both θ (in the x-z plane) and ϕ (in the y-z plane). If the incidence beam has an angular spread, the ensemble of the shift effectively blurs k_T and reduces the resolution of the spectrometer. Due to the asymmetry of the wave vector deflection introduced by the grating, the effective resolution of the spectrometer suffers differently for θ spread and ϕ spread. In Figure 3 (c), we plot the theoretical effective resolution over incidence angle spread as a function of θ spread and ϕ spread under a 30° imager tilt, where the resolution under normal incidence is below 0.5 nm at $\lambda = 830$ nm. The plot suggests that in order to maintain the resolution around one nanometer in this ideal model, one needs to restrain the angular spread θ to be within around one degree and ϕ to be within around one hundredths of a degree. This asymmetry of the dependence of effective resolution over θ spread and ϕ spread for the Talbot spectrometer is similar to that of a conventional spectrometer, where the spectrometer resolution is more sensitive over the width direction of the slit than the length direction of the slit due to the asymmetric dispersion from 1D gratings.



Fig. 3. (a) Wave vector scattering due to grating under normal and oblique incidence. (b) Illustration of the incidence angle spread on the grating surface used in calculating the plot in (c). (c) Effective resolution as a function of θ spread and ϕ spread for a camera tilt angle of 30°.

5. Conclusions

In conclusion, our results demonstrate the potential of utilizing mid-field Talbot effect under non-paraxial diffraction for building compact spectrometers. Our model agrees well with experiment and can serve as a simple design guideline for constructing spectrometers with a given target performance metric.

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